

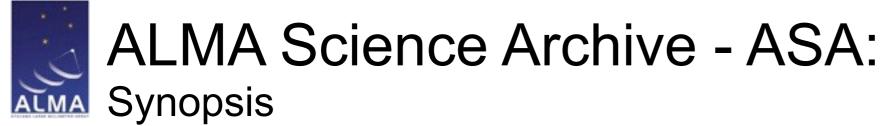




Using ALMA Archive Substructure and Operation

Erik Muller, Ken'ichi Tatematsu, Daisuke Iono.

EA-ARC ALMA user support (NAOJ Chile Observatory)





- 1. ASA: Motivated by Scientists
- 2. Structure of the Archive facility
- 3. Obtaining data from the archive
- 4. Structure of Archived data.
- 5. PI publication requirements
- 6. Future development.





What do scientists want?

- 1. Reduced, science-ready, high-level data (PIs want proprietary protection)
- 2. Processible, accessible data
- 3. Data they can reprocess/analyse with their "pet" data analysis software

Solution:

- 1. Provide processed, calibrated data, with clear description of the processing
- 2. Provide general format data
- 3. Provide raw data for the PI to *completely* re-process, if desired.

ALMA science Archive caters for TWO "clients":





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ALMA science Archive caters for TWO "clients":

 The principle investigator and "delegated" collaborators.
 (Exclusive access during the proprietary period).

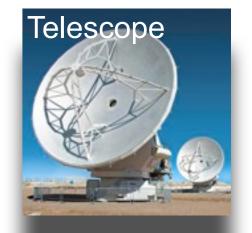
 Future Archive Researchers
 (After Proprietary period expires, data is accessible by anyone).



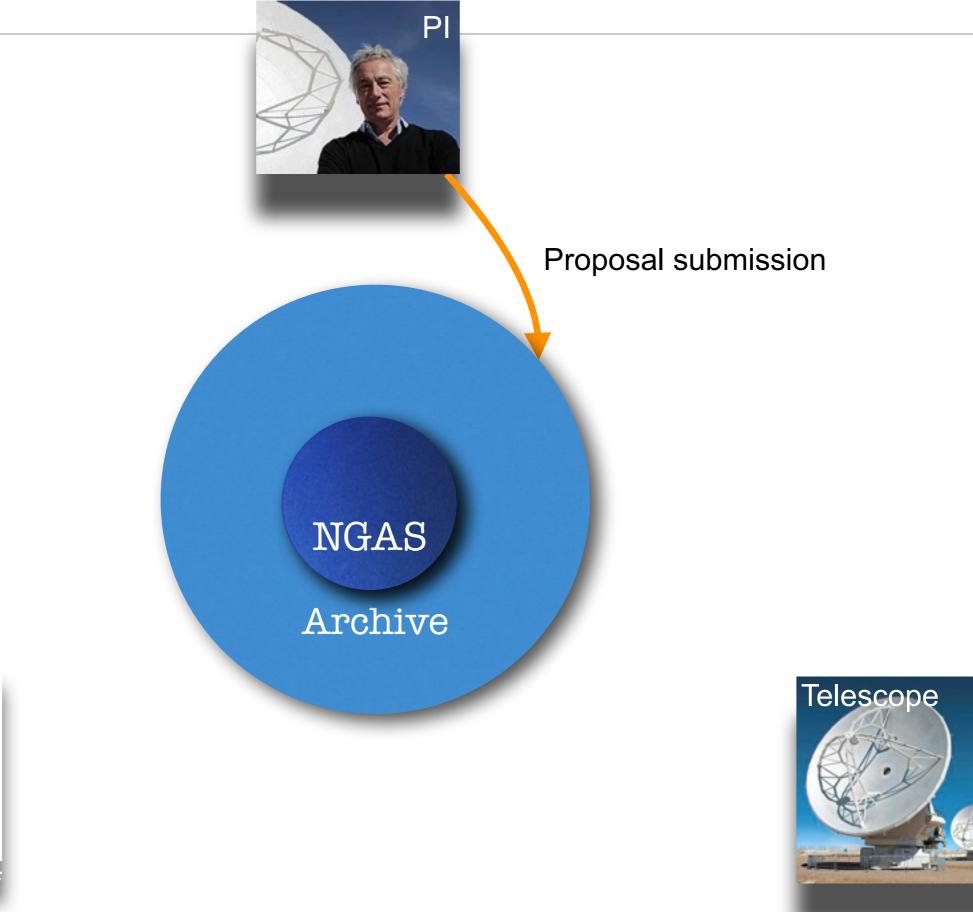


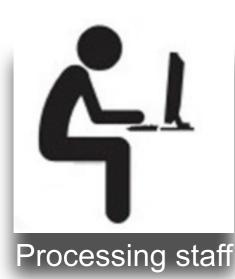




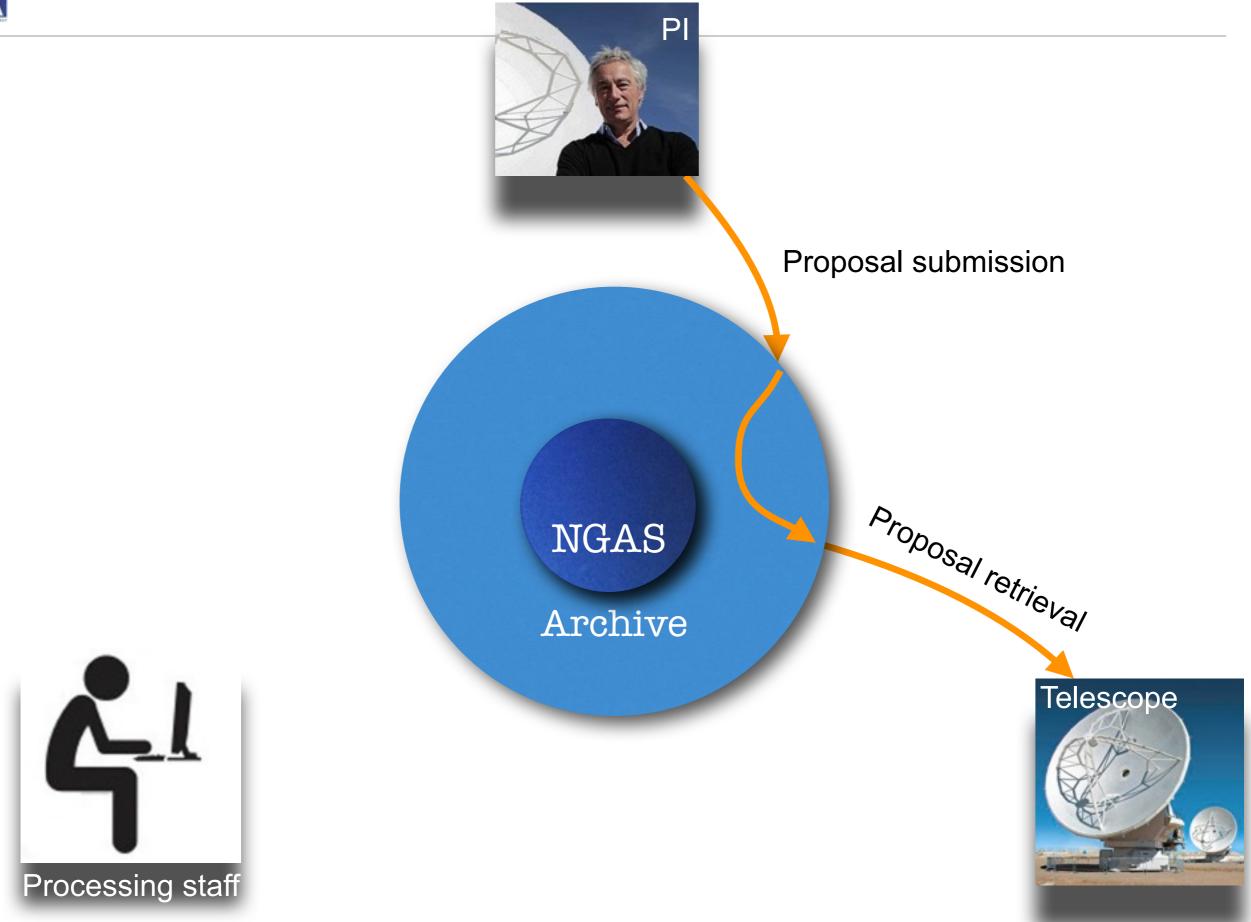




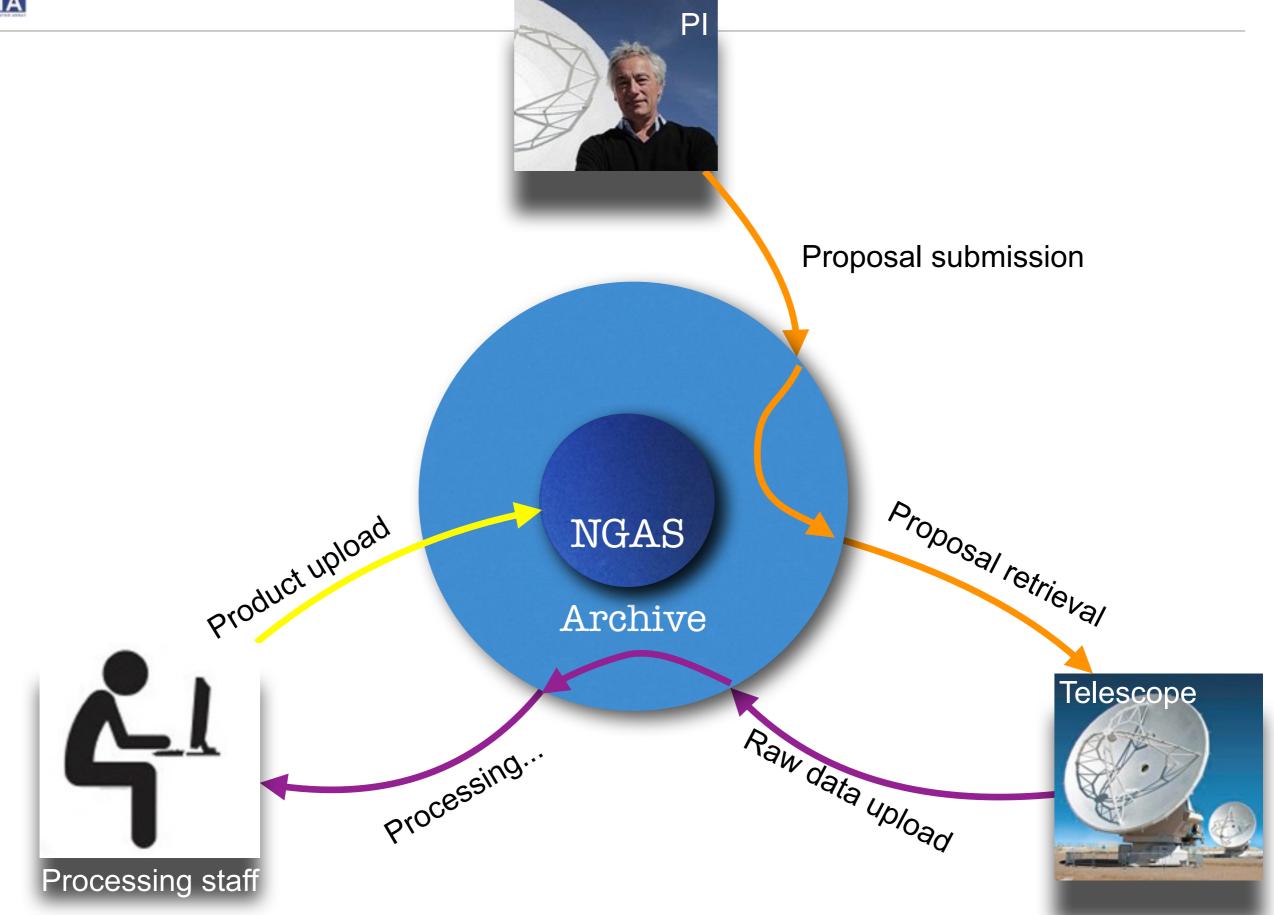


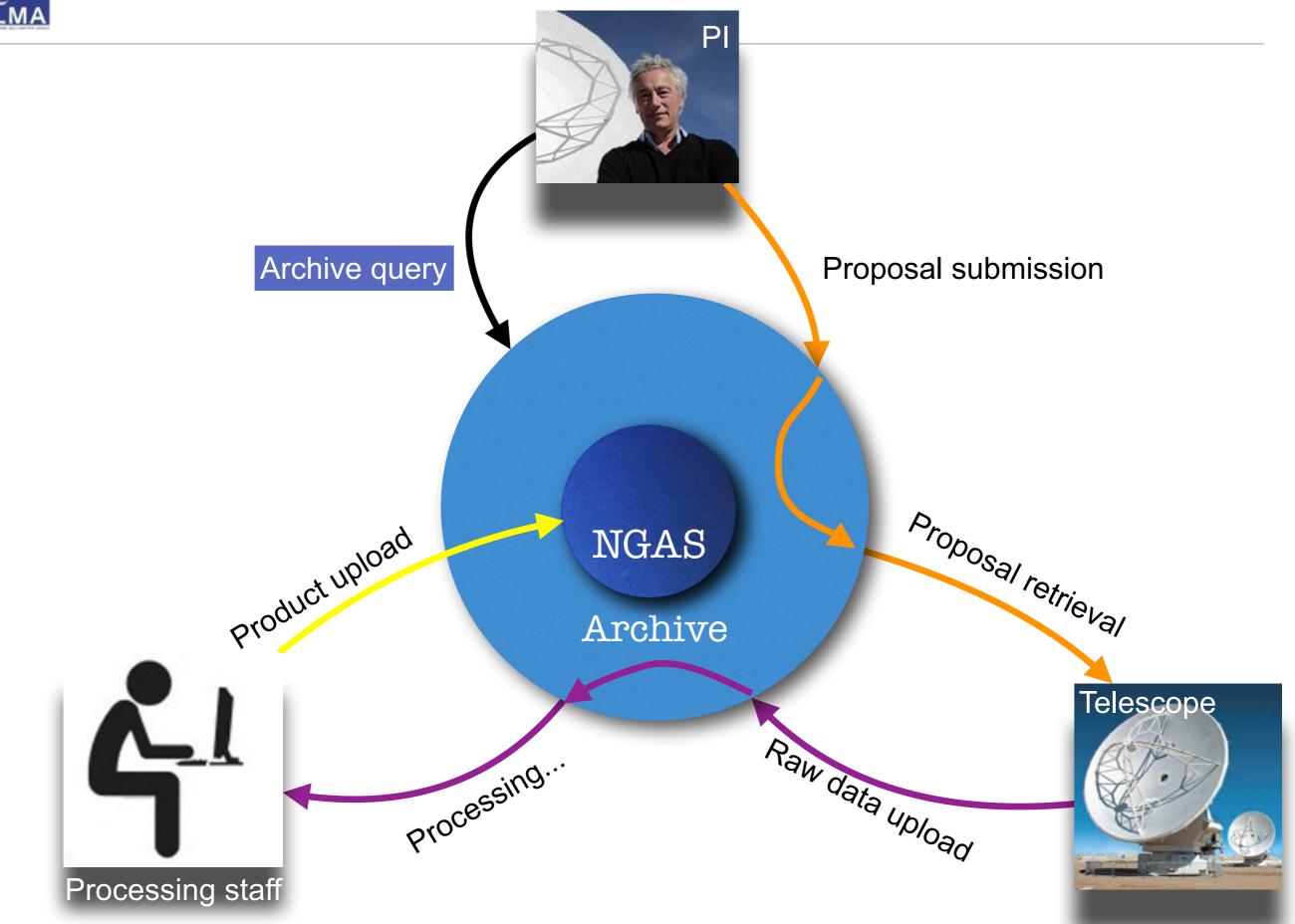


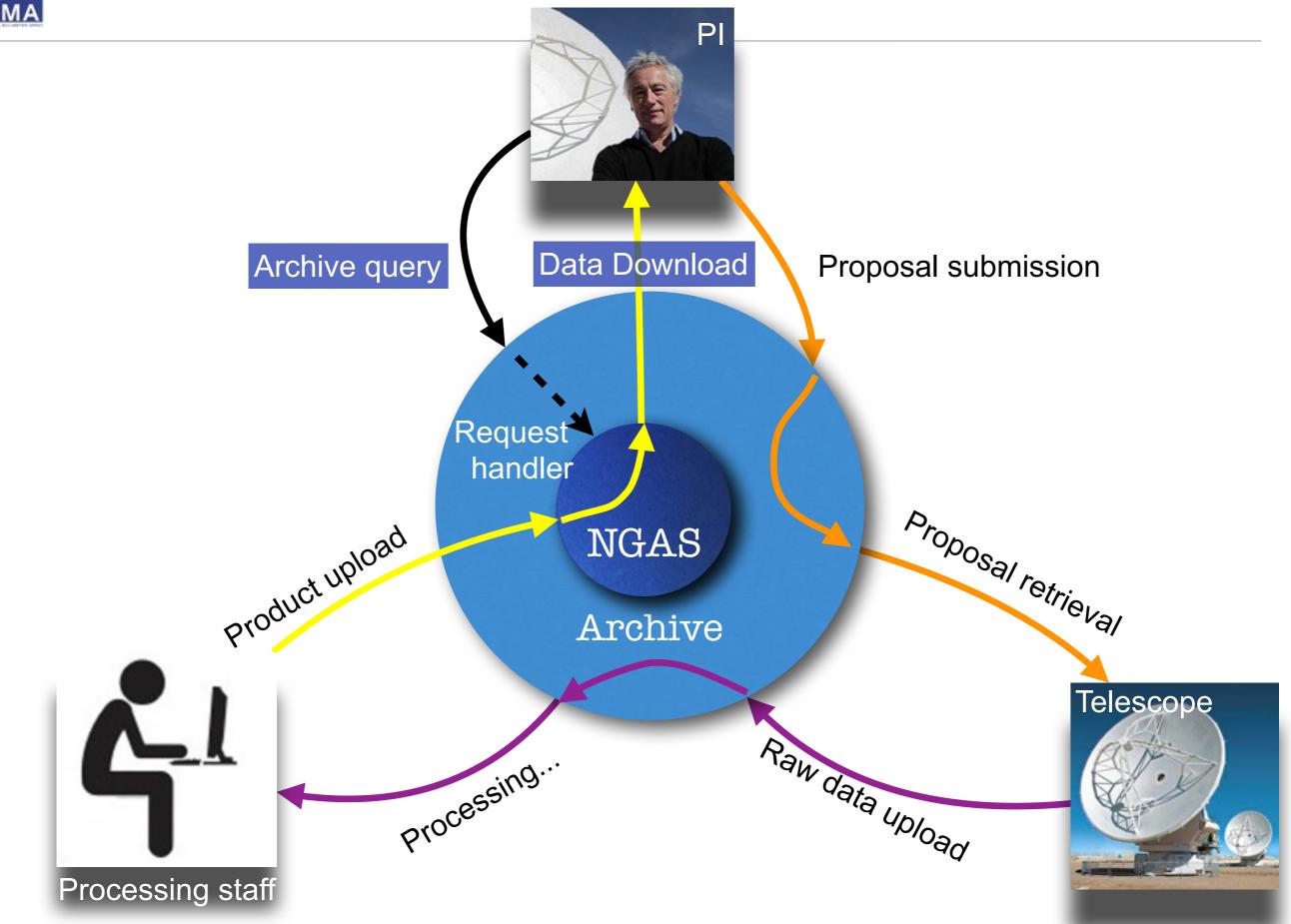
















Current status

- Presently, EA is delivering data primarily by FTP transfer.
 (Resulting in a shorter delivery time, after staff-processing stage is complete)
- We intend to migrate to delivery by Archive internet data distribution, in the near future (~3-5 weeks).
- Some deliveries are already being made using Archive internet data-distribution.
- Delivery is currently effected through one or more 'tarballs'.
- Raw data ('ASDMs') not 'delivered' by default in cycle 1+ data, but are available directly by internet request.
- Future researchers can optimise pipeline processing for their science





PI notice

- PI is currently notified by email, using ALMA helpdesk.
- Email will contain instructions on how to obtain your data, one of two methods:
 - 1. by FTP (File transfer Protocol), from EA-ARC firewall-protected repository
 - 2. from the Archive data distribution system, using your ALMA login.

PI - Notification for Delivery by FTP transfer

Dear Albert Einstein, We are pleased to announce that data of the Scheduling Blocks "alpha_cen_12m" containing execution block uid___A00_X123_X456 from your project 2012.1.00111.S with the title "Alpha centauri and all that is beautiful" are now available to you for download. A tarball of these data can be obtained from (1) Using web browser: FTP URL: ftp:// (2) Using FTP-specific client software: Host: alma-dl.mtk.nao.ac.jp User: [user] Path: [path] Only passive-mode FTP is available. The password will be sent to you via separate email. Please reply to this ticket if you do not receive a password within 1hour.

The account will expire at 19th april,16:00 JST





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"Alpha centauri and all that is beautiful" are now available to you for download. A tarball of these data can be obtained from the ALMA Science Archive.
Please use your ALMA userid "username" to login to the ALMA Science Portal, http://almascience.nao.ac.jp
and follow the "Data" -> " link to the archive to search for and download your data, or access to
http://almascience.nao.ac.jp/aq/asdm_uid=uidA00_X123_X456
A download request of both processed and raw data can be submitted by selecting all data (topmost checkbox) and hitting "submit download request".

ALMA Science Archive - ASA: Delivery: PI and AR - what do you get?

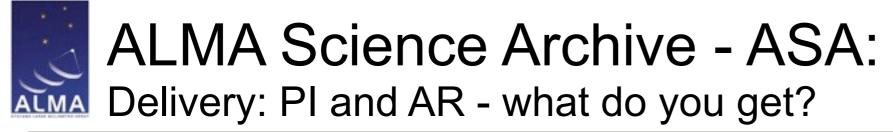
Delivery content for Cycle 1 deliveries + onwards.

- FITS LOGS Scripts Cal. Tables QA rept. README
- Data products (FITS cube)
- Observing Logs
- Calibration, processing + Imaging Scripts
- Calibration tables
- Quality Assesment (QA0, QA2) Reports
- README file (text file describing contents of delivery)











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- Data products (FITS cube)
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- Calibration, processing + Imaging Scripts
- Calibration tables
- Quality Assesment (QA0, QA2) Reports
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Delivery to PI via network





Proprietary rules:

- PI Data is available for public use after 12 calendar months (6 months for director discretionary time)
- Science verification data available after verification

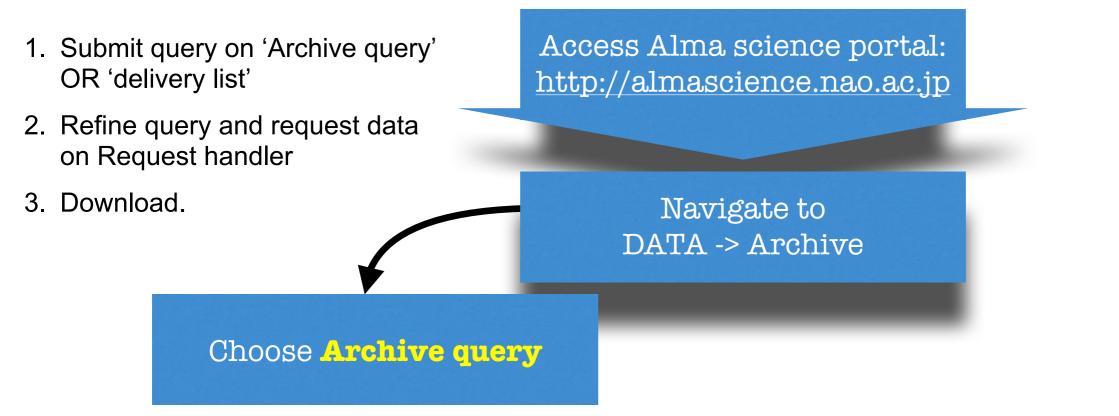
	Accessible only by PI	Accessible to anonymous
ALMA PI & delegated COIs	QA2 12 Mnths after delivery	
Science verification	QA2 verification	

Proprietary-expired data are all available through the Science portal.

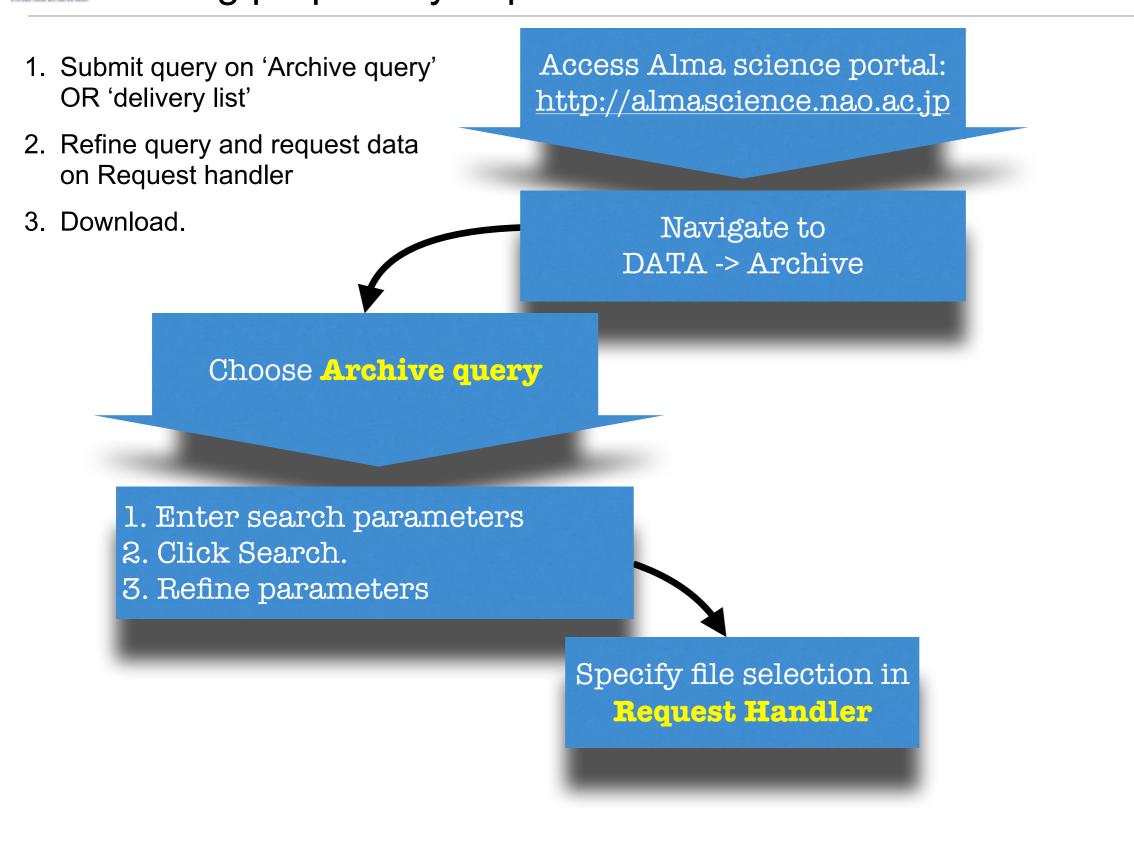


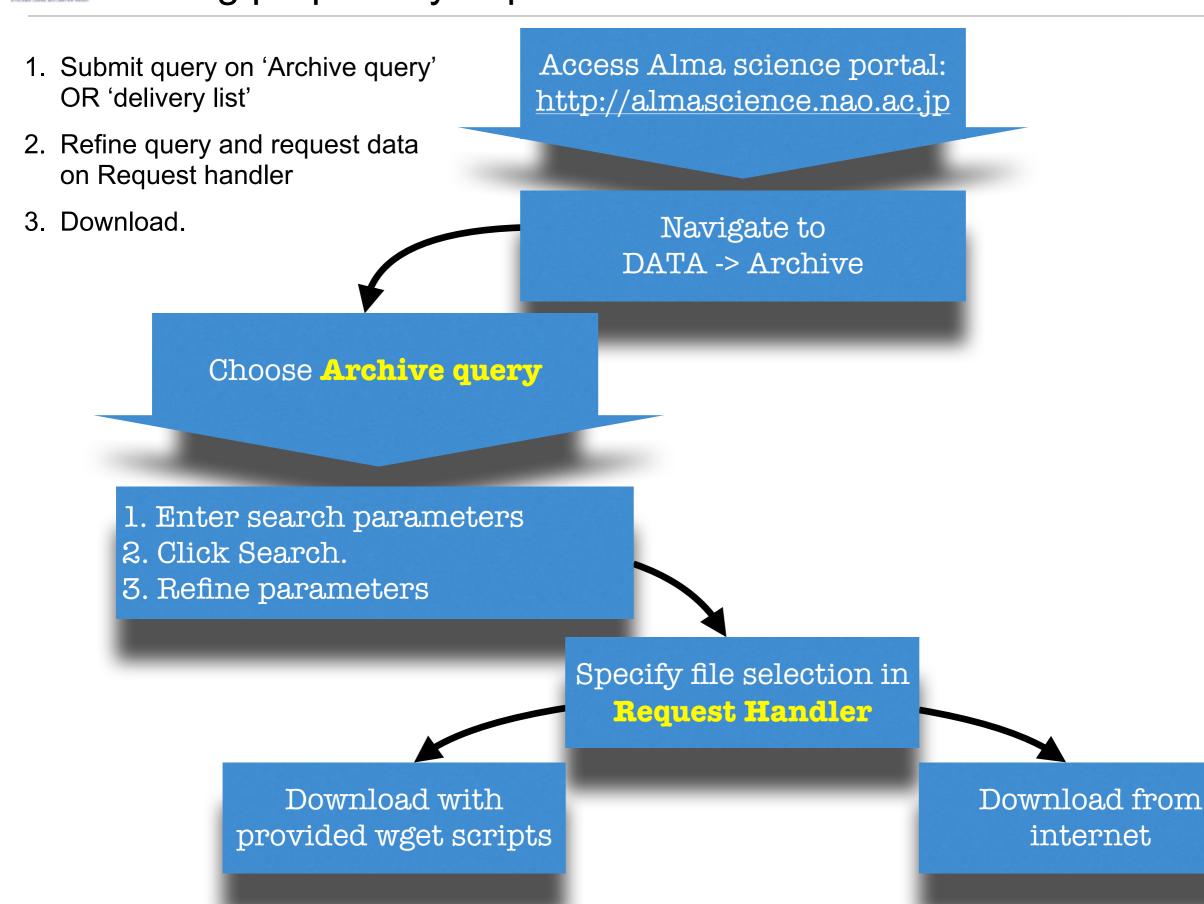
Getting proprietary-expired data:

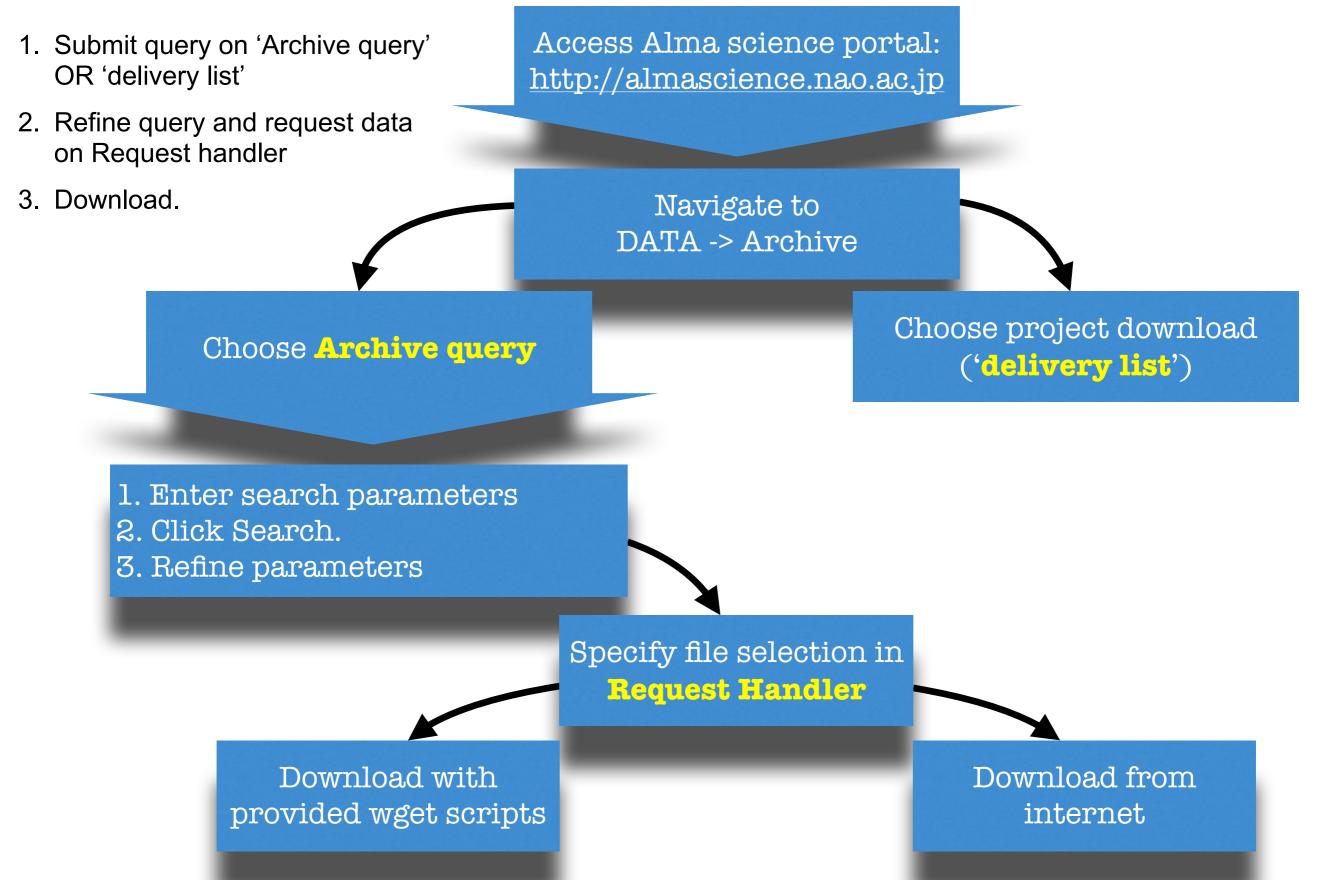
- 1. Submit query on 'Archive query' OR 'delivery list'
- 2. Refine query and request data on Request handler
- 3. Download.

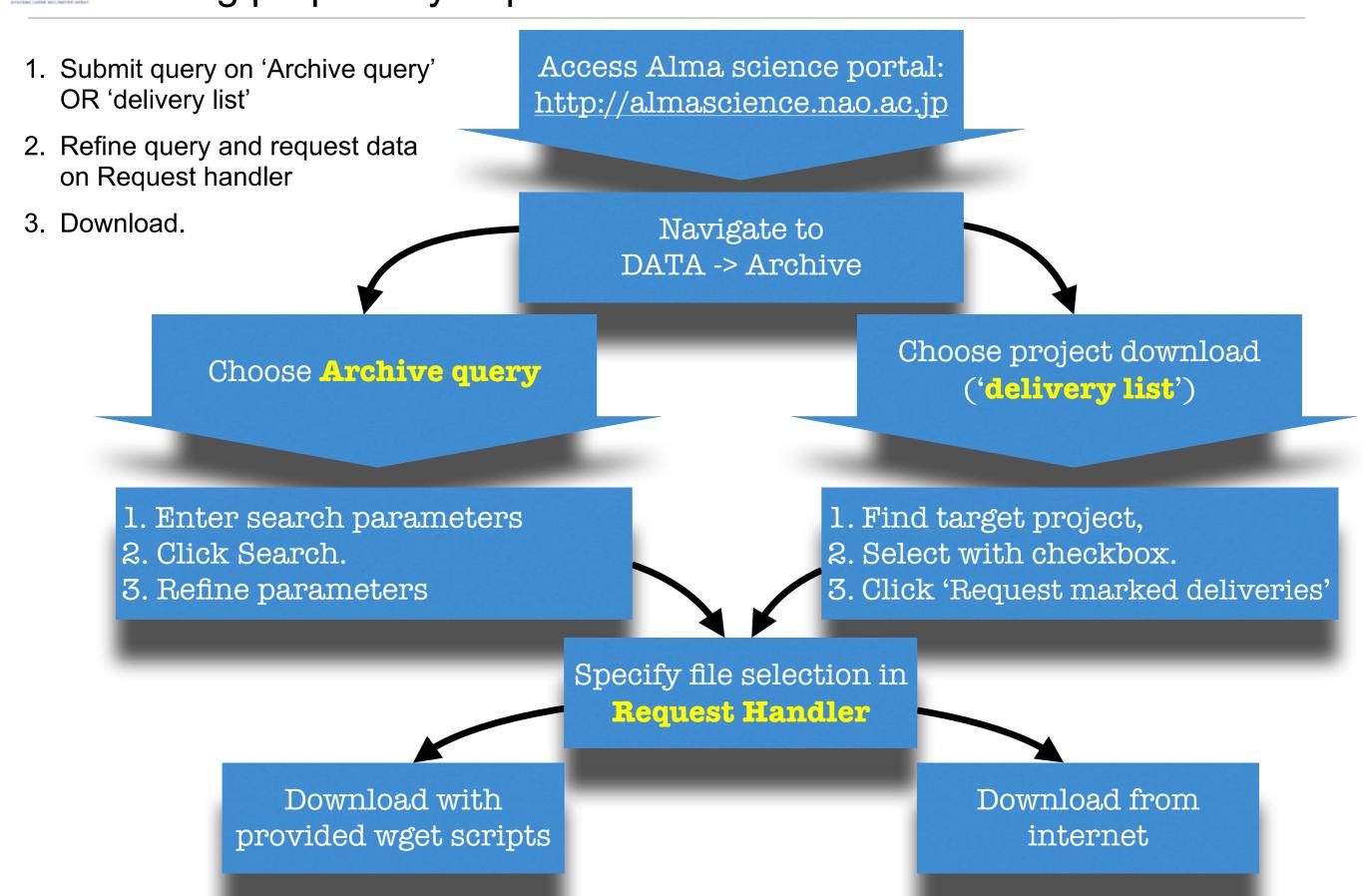


Access Alma science portal: 1. Submit query on 'Archive query' OR 'delivery list' http://almascience.nao.ac.jp 2. Refine query and request data on Request handler 3. Download. Navigate to DATA -> Archive Choose Archive guery 1. Enter search parameters 2. Click Search. 3. Refine parameters









Archive Query:

When using JAVA 7.51+: security settings should be modified.

Atacama Large Millimeter/Submillimeter Array In search of our Cosmic Origins								
You are here: <u>Home</u> > <u>ALMA Data</u> > Archive ALMA Science Ar Query Form Result Table Search Reset								
Position Source name (Sesame) Source name (ALMA) RA Dec Search radius 0:10:00	Energy Frequency Bandwidth Spectral resolution Band	Time Observation date Integration time	Polarisation Polarisation type					
Observation Water vapour Scan intent Observe target	Project code Project Title PI Name		Options Results view raw data project					

- Enables 'standard' search patterns
 - target, position, band. Observing date & conditions. Also: project code, PI name
- Enables advanced filtering capbilities
 - Limit search by linking logical filter parameters (band, sensitivity, etc. etc.)
- Download links are returned to user.



Request Handler

	Atacama Large Millimeter/Sub In search of our Cosmic Origins	omillimeter Array		
	tequest Handler			Login
Archiv	ve Requests Req #85,290,726			
		Select All Unselect model of complete request: downloadReguest85290726script.sh		d Selected
UNIX/Lir	nux Shell script if you prefer command line down	nload of complete request: downloadRequest85290726script.sh		
	nux Shell script if you prefer command line down Data en Project / OUSet / Executionblock	load of complete request: downloadRequest85290726script.sh		Access
JNIX/Lir	nux Shell script if you prefer command line down	nload of complete request: downloadRequest85290726script.sh		
JNIX/Lir	nux Shell script if you prefer command line down Data en Project / OUSet / Executionblock	nload of complete request: <u>downloadRequest85290726script.sh</u> Intities 1-1 of 1 File	Size	
Select	nux Shell script if you prefer command line down Data en Project / OUSet / Executionblock	nload of complete request: <u>downloadRequest85290726script.sh</u> Intities 1-1 of 1 File 2011.0.00099.S_2012-02-24_001_of_004.tar	Size 19.1GB	Access
Select	nux Shell script if you prefer command line down Data en Project / OUSet / Executionblock	nload of complete request: <u>downloadRequest85290726script.sh</u> <i>ntities 1-1 of 1</i> File <u>2011.0.00099.S_2012-02-24_001_of_004.tar</u> <u>2011.0.00099.S_2012-02-24_002_of_004.tar</u>	Size 19.1GB 16.8GB	Access

- Enables selection of specific tarballs (note, duplicity exists in some cases, for cycle 0 data)
- Provides a text linux script invoking **wget** commands, which may be run from the terminal.



ALMA

ALMA Science Archive - Archive research NACJ Cycle 0 - vs - cycle 1

- Cycle 0 data includes the raw data. (in an ASDM format file e.g. uis___X123_X123_X12.asdm.sdm.tar)
- Cycle 0 data also contains the 'measurment set' data, which is calibrated
- Cycle 0 data also contains the products in FITS format
- Cycle 1 data tarballs to NOT include the raw data, nor the calibrated measurement set.
- Cycle 1 data tarballs include the products, in FITS format.

Request #85314508 by Anonymous User <

NIX/Linux Shell script if you prefer command line download of complete request: downloadRequest85314508script.sh							
Data entities 1-1 of 1							
Select	Project / OUSet / Executionblock	File	Size	Access			
	uidA001_Xa0_X17e5						
		2012.1.00941.S_uidA002_X5d7935_X1cb_001_of_001.tar	85.9MB	0			
		2012.1.00941.S_uidA002_X5d7935_X1db_001_of_001.tar	243.8MB	0			
		uidA002_X67f61e_X3a4.asdm.sdm.tar	1.2GB	0			
		uidA002_X67f61e_X4c2.asdm.sdm.tar	1.2GB	0			
		uidA002_X67f61e_X57a.asdm.sdm.tar	1.2GB	0			
		uidA002_X6a1245_X976.asdm.sdm.tar	1.2GB	0			
		uidA002_X7330a2_X11b5.asdm.sdm.tar	1.9GB	0			
	Dat	a entities 1-1 of 1	6.9GB				

ALMA Science Archive - Archive research What next? for PIs and Archive researchers.

- You will receive one, or many tarballs untar these (macs can't do this properly from a on-screen click. Untar the folders from the command line)
- Examine the README file, it contains information about the contents of the tarballs
- Examine the PRODUCTS directory, it contains FITS-format images of the target frequencies (or continuum), made by the PI

Atacama Large Millimeter/submillimeter Array (ALMA)

#####

Cycle: O (Early science) Project code: 2011.0.00199.SProject title: Bursting Water Maser Feature in Orion KL Configuration: Extended Proposed rms: 170mJy CASA version used for reduction: 3.4 Comments from Reducer:

DA42, DV01, DV07, and DV23 were removed because of large pointing error. DV17 was removed due to negative Trec. DV19 XX pol. was removed due to out of spec Trec temperature and high Tsys. DA51 was removed due to bad bandpass gain solutions. DV05 was removed due to poor S/N in bandpass.

The rms was achieved to be 105 mJy (the required rms is 140 mJy) by the integration in 7.5 GHz in frequency width but the image may contain non-real det ection.

WVR correction seems to make phase worse in some antennas. It is probably better not to use them in those antennas. emuller@alma-work:~/2011.0.00199.S\$ more `find -name README` Atacama Large Millimeter/submillimeter Array (ALMA) (C0, C1+) README contains

- 1. Basic information about the data.
- 2. Notes from the data reducer: flagging, high tsys, etc.
- 3. A directory tree

ALMA Science Archive - Archive research What next? for PIs and Archive researchers.

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- 2011.0.00199.S/ -- sg_ouss_id/ -- group_ouss_id/ -- member_ouss_2013-03-01_id/ -- README -- product/ _A002_X590461_X620.ms.split.image.continuum.source4.image.fits -- uid _A002_X590461_X620.ms.split.image.continuum.source4.mask/ -- uid_ -- calibration/ -- uid____A002_X590461_X620.calibration.plots/ _A002_X590461_X620.calibration/ -- uid -- log/ __A002_X590461_X620.calibration.log l-- uid -- raw/ -- uid____A002_X590461_X620.ms.split/ -- calibrated/ -- uid____A002_X590461_X620.ms.split.cal/ -- qa/ A002_X590461_X620__qa2_part1.png -- uid _A002_X590461_X620__qa2_part2.png -- uid __A002_X590461_X620__textfile.txt -- uid __A002_X590461_X620__qa2_part3.png -- uid -- script/ _A002_X590461_X620.ms.scriptForCalibration.py - uid _A002_X590461_X620_scriptForImaging.py -- uid

(C0, C1+) README contains

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- 3. A directory tree

ALMA Science Archive - Archive research Raw data, and measurement sets in Cycle 1+

How to make cycle 1 measurement sets (refer to README information):

- Cycle 1 data includes the CASA processing script, which runs on the raw data (asdm format)
- PIs and Archive users can obtain the raw data from the archive, put it in the 'calibrated' directory, and run the script (or adjust it)

Why do we do this?

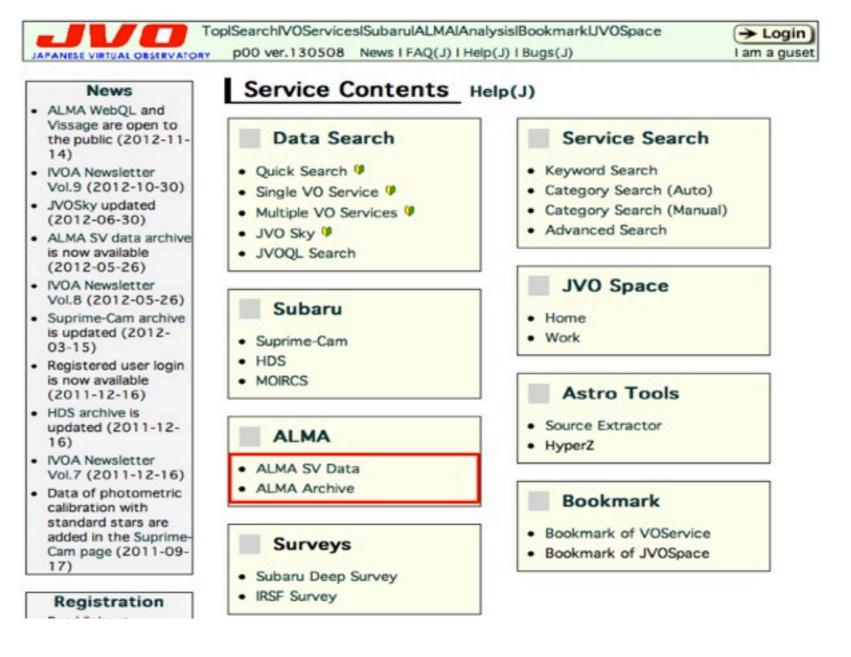
- Not all PIs/ARs need the measurement set, it is very bulky (~100's GB), and slows transfer
- Not all PIs/ARs need the raw data, but it must be available.

ALMA Science Archive & JVO:



http://alma-intweb.mtk.nao.ac.jp/~kt/ALMA_archive_handbook.pdf

locally-stored data query interface



ALMA Archive : Target Info

Target Name: R Scl

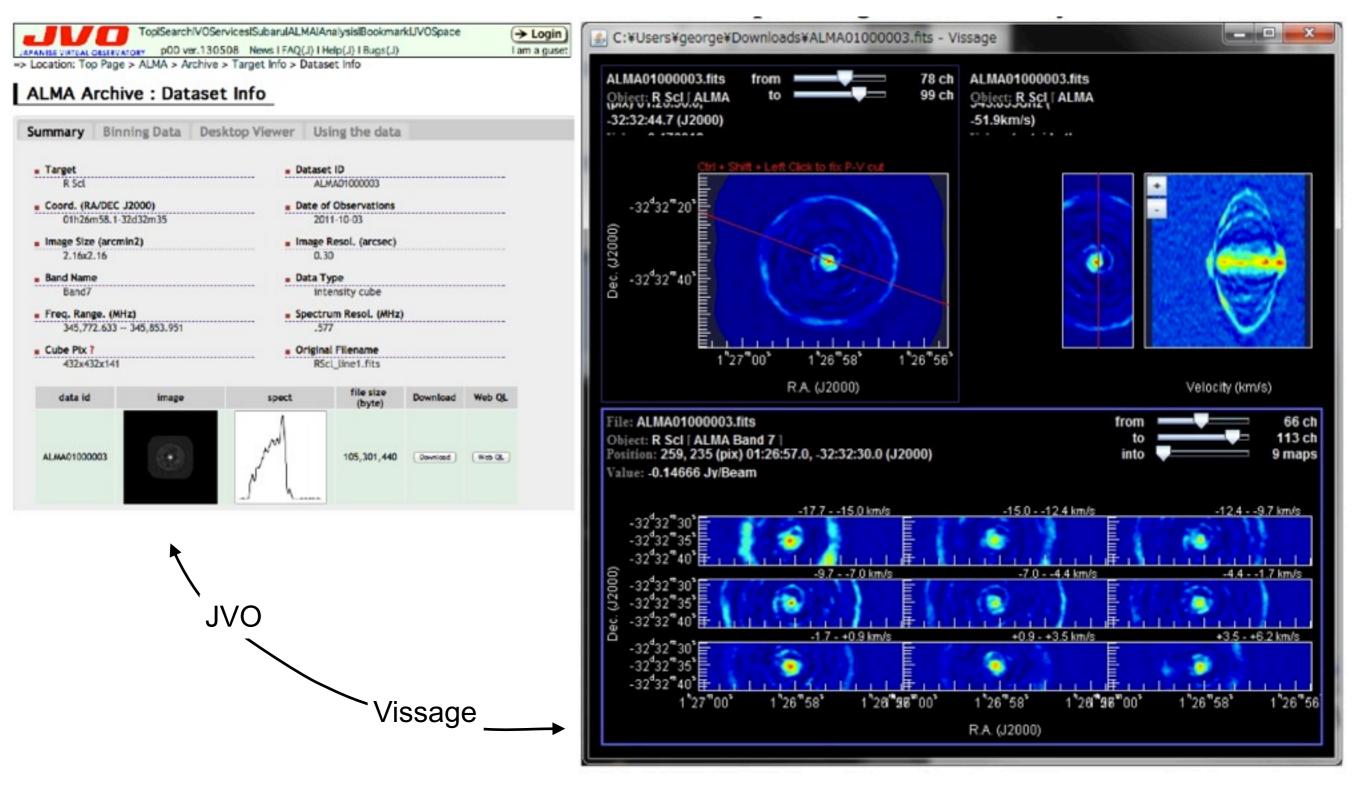
#	dataset id	ra/dec (J2000)	size (arcmin2)	band	freq. range (GHz)	data type	Cube size (XxYxF) ?	image resol (arcsec)	freq. resol (MHz)	obs date	original fits name
1	ALMA0100002	01h26m58.1-32d32m35	2.16x2.16	Band7	330.232 - 333.962	intensity cube	432x432x1	0.30	3,729.650	2011-10-03	RScl_continuum.fits
2	ALMA01000003	01h26m58.1-32d32m35	2.16x2.16	Band7	345.773 - 345.854	intensity cube	432x432x141	0.30	.577	2011-10-03	RScl_line1.fits
3	ALMA01000004	01h26m58.1-32d32m35	2.16x2.16	Band7	345.328 - 345.398	intensity cube	432x432x121	0.30	.576	2011-10-03	RScl_line2.fits
4	ALMA01000005	01h26m58.1-32d32m35	2.16x2.16	Band7	342.871 - 342.940	intensity cube	432x432x121	0.30	.572	2011-10-03	RScl_line3.fits

ALMA Science Archive & JVO:



http://alma-intweb.mtk.nao.ac.jp/~kt/ALMA_archive_handbook.pdf

Purpose-built data visualisation with "Vissage"





ALMA Science Archive - ASA:



Required acknowledgement information.

Please notify EA-ARC and EA-ARC outreach for any publications making use of ALMA data.

Publications making use of ALMA data (PI and Archive researchers) are required to include the following acknowledgements in publications:

"This paper makes use of the following ALMA data: ADS/ JAO.ALMA#2011.0.01234.S . ALMA is a partnership of ESO (representing its member states), NSF (USA) and NINS (Japan), together with NRC (Canada) and NSC and ASIAA (Taiwan), in cooperation with the Republic of Chile. The Joint ALMA Observatory is operated by ESO, AUI/NRAO and NAOJ."



ALMA Science Archive - ASA:



Where to now?

Immediate Future improvements

- 1. Basic layout to be modified slightly
- 2. 'Frequency support' information to be improved
- 3. File Download destination support improved
- 4. Filtering improved
- 5. online help improved.
- 6. Visual quick-look planned.

Not happy with the Archive functionality?

Or are you having problems?

YOU are the users, but we need your input to make the system better for everyone. We can only improve the archive if you tell us about problems, or any suggestions.

If you have any comments or suggestions, please submit them to EA-ARC using the helpdesk system (https://help.almascience.org).

Please also ensure you examine the 'help' information on both the Archive query and Request handler pages.

Please contact EA-ARC for face-to-face support information